



Australian Government Geoscience Australia

Positioning Australia

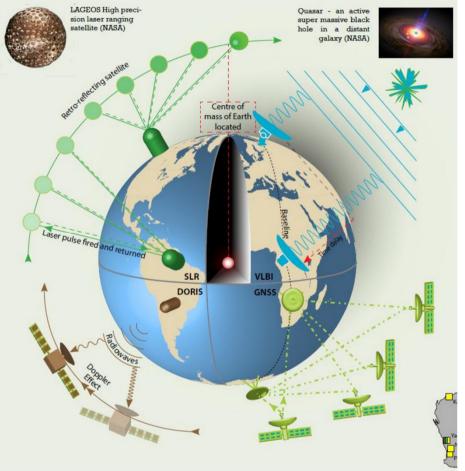
# **Balancing Best Practices** with Innovation:

Developing a sustainable program for Local Tie Surveys at Australian Geodetic Observatories

Bart Thomas | Geodetic Surveyor

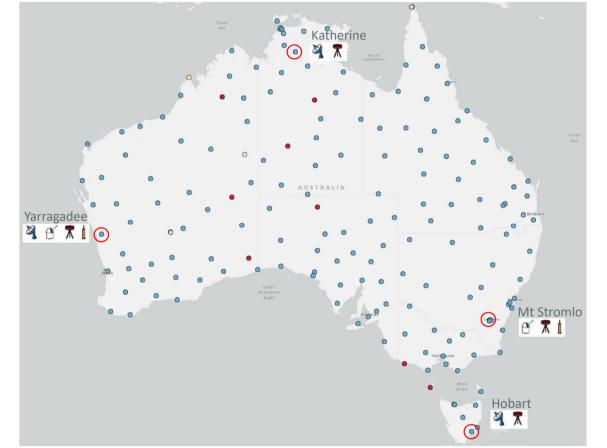
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## Local Tie Survey of Co-located Systems



- The International Terrestrial Reference Frame (ITRF) Realisations are the combination of the four space geodetic techniques that are part of the Global Geodetic Supply Chain
  - Satellite Laser Ranging (SLR)
  - Global Navigation Satellite Systems (GNSS)
  - Doppler Orbitography Radio-positioning by Satellite (DORIS)
  - Very Long Baseline Interferometry (VLBI)
  - Co-locating these techniques allows for their combination

No matter how accurate the individual techniques are, a ground survey ties them all together



### **Domestic Co-Located Geodetic Observatories**

TOTAL: 20 CO-LOCATED FEATURES

#### Yarragadee, WA

- 12m VLBI Radio Telescope (AuScope)
- Satellite Laser Ranging Telescope
- GNSS (x 3)
- DORIS
- Gravity BM

#### <u>Mt Stromlo, ACT</u>

- Satellite Laser Ranging Telescope
- GNSS (x 3)
- DORIS
- Gravity BM

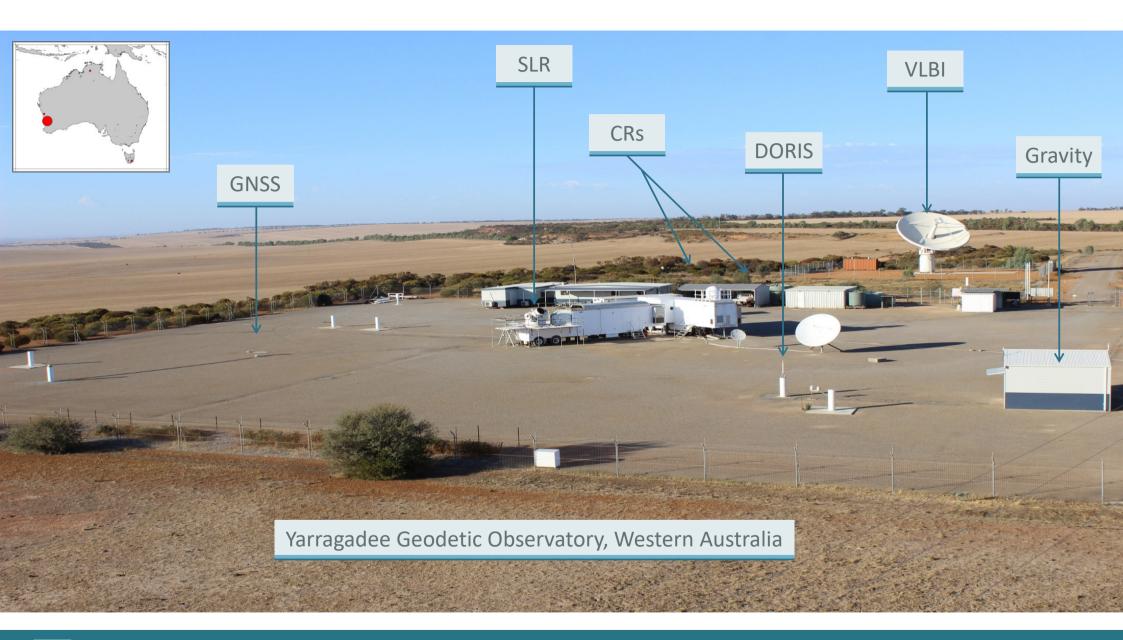
#### Katherine, NT

- 12m VLBI Radio Telescope (AuScope)
- GNSS (x 2)

#### <u>Hobart, TAS</u>

- 12m VLBI Radio Telescope (AuScope)
- 26m VLBI Radio Telescope
- GNSS (x 1)
- Gravity BM

(https://gnss.ga.gov.au/)



## **GA Local Tie program**

- 2023
  - Re-development of the capacity & plan to provide an updated set of results
  - Re-engagement with the international and educational communities, and local stakeholders of each site.
  - Review/modernisation of the technique to align with available resourcing, hardware, and software limitations, for sustainable continuation of the program
  - Maintenance and Calibration of equipment from the AuScope Australian Geophysical Observing System (AGOS)<sup>1</sup> pool of Geospatial Equipment, required to complete the survey
  - Publication of recent outstanding reports (Mt Stromlo 2014, & 2018)
  - Observation of Hobart Site and test / re-discover the established process
- 2024
  - Observation of **Mt Stromlo, Yarragadee & Katherine**, and further refinement of the process
  - Improved balance of WHS considerations in the observation process.
  - Refinement of the process and establish the future direction of the program
  - Publication of 2024 Reports

<sup>1</sup>https://www.auscope.org.au/geodesy





A Quick Summary of Why and How...

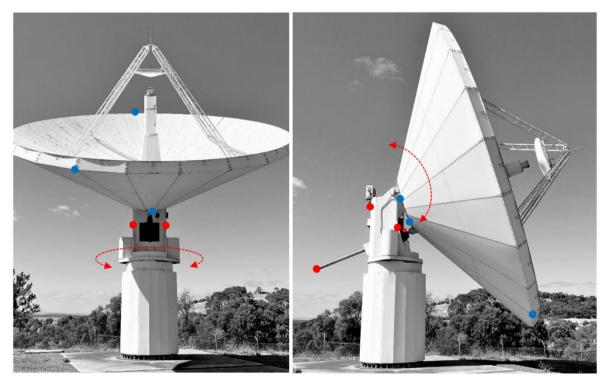
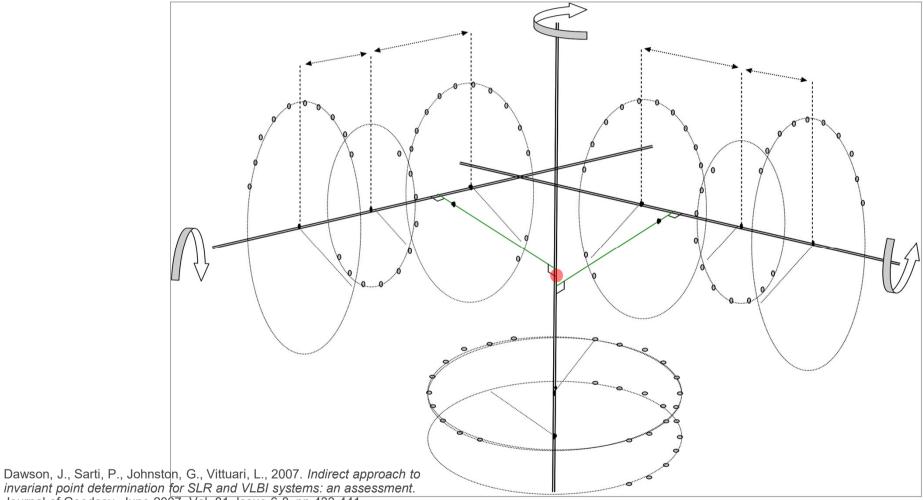


Figure 3.2.2.1 Primary Azimuth Axis

Figure 3.2.2.2 Secondary Elevation Axis



## A Quick Summary of Why and How...

*invariant point determination for SLR and VLBI systems: an assessment.* Journal of Geodesy. June 2007, Vol. 81, Issue 6-8. pp 433-441.

## **The Established Survey Method:**

- 1. Full survey network observation
  - 1. Orientation
  - 2. Levelling survey
  - 3. Horizontal angles
- 2. Most accurate equipment
- 3. Best practice measurements
  - ICSM SP1 Guideline for Conventional Traverse Surveys v2.2
- 4. High number of observations on telescope (rotation and targets)
  - 1. Time and Safety

Prism	Constant [mm]	Centring Accuracy [mm]	Dimensions [mr
GPH1P ular Prism	0.0	0.3	118 51 85 R 39.5

What Can we improve?

Time, Safety, Complexity, Consistency?

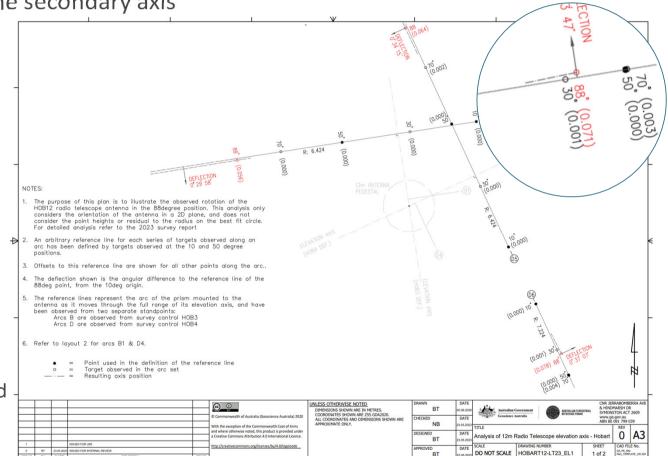
## **Discovering what we don't know:**

#### Detecting movement in the motion of the secondary axis

- Solution showing large residuals in circle fit
- Further analysis showed the issue was in the higher elevation obs.
- Mechanical failure was ruled out

#### Outcome:

- The source was determined to be a calibration model applied to the pointing of the antenna
- The effect was observed in unresolved data from a previous survey
- Identified a need for better recording of the survey methodology
- The method to move the antenna was changed to 'maintenance mode'
  - Development of SOP's.

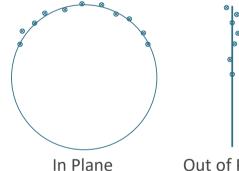


## **Refining the Survey Method:**

#### SAFETY:

- Use of 360 prism to reduce re-pointing and the number of times elevated to height
  - Over a working distance of <50m
  - RMS & Mean residual to circle fit are indistinguishable.
- Use of light weight Tooling Ball prisms to remove personnel from SLR telescope altogether
- Both have benefits of SAFETY & TIME improvement
  - Faster Setup & observation sequence
  - Less manual pointing
  - More data points observed as a result

Prism	Constant [mm]	Centring Accuracy [mm]	Dimensions [mm]	1	
GMP101 MiniPrism	+17.5	1.0	25 <b>15</b> 138	1	
GRZ101 360° MiniPrism	+30.0	1.5	39		



05

0.6

0.1

W1 W2 W3 Y1 Y2 Y3 Average

0.7 0.3 0.6 0.6 0.4 0.4

0.5 0.8 0.5 0.5 0.8 0.6

-0.3 -0.2 0.3 0.2 0.2 0.0

Average In Plane RMS Difference

0.2 0.1 0.3 0.0 0.2 0.2

A7 A8 A9 C7 C8 C9

Azimuth Axis Circle Fit Residuals (mm) GMP101 Prism/Circle

In Plane Mean

Prism/Circle

In Plane RMS

In Plane Mean

(Yarragadee) In Plane RMS

**GRZ101** 

(Katherine)

#### Out of Plane

Azimuth Axis	s Circle Fit Residuals (mm)							
GMP101	Prism/Circle	W1	W2	W3	Y1	Y2	Y3	Average
Yarragadee	Out of Plane RMS	0.2	0.2	0.5	0.2	0.2	0.2	0.2
	Out of Plane Mean	0.0	0.0	0.0	0.0	0.0	0.0	
GRZ101	Prism/Circle	A7	A8	A9	C7	C8	C9	
Katherine	Out of Plane RMS	0.8	1.1	0.8	1.0	0.9	0.9	0.9
	Out of Plane Mean	0.0	-0.1	0.0	0.0	0.0	0.0	
		Average C	Out of	Plane	RMS	Diffe	rence	0.7

## **Reducing Complexity - The Mt Stromlo Local Example**

#### Assumptions:

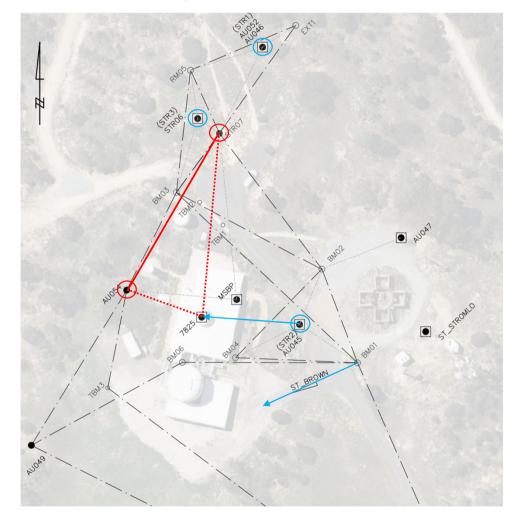
- Survey control is rigid (only observe from established control)
- GNSS CORS are on stable monuments, self monitoring & have been omitted from this survey.
- Absolute position and orientation were previously defined.

#### <u>Outcome</u>

- Deformation monitoring for the SLR telescope
- Develop a safe and adequate methodology
  - Protect the personnel by removing during observation
  - Protect the SLR equipment
- Compare to the more rigorous result obtained in 2024

#### Local Topocentric Coordinate System

	Origin	IVP		^е	^n	^u
2024	AU45	:	7825	-24.9525	2.0070	2.4953
2025	AU45	:	7825	-24.9536	2.0069	2.4962
				-0.0011	-0.0001	0.0009



## **Conclusion & Future direction of local tie surveys by Geoscience Australia**

Adopting these findings:

- Local Tie Surveys can be completed more consistently with less resources and time required
  - Leading to a more sustainable work program and regular results
- Major effort in the full site survey can be reduced in frequency or 'as required' (ITRF contribution)
  - Validation surveys could be completed between these epochs
- Simplified methodology could reduce effort further by;
  - Training local staff at the observatories to complete the measurements
  - Contracting the work out to local Surveyors.
  - Installing semi-permanent targets, mounts or equipment
- Development of a Point Cloud Observation Methodology
  - Potentially could derive a solution for telescope IVP from normal operation
  - Step towards automation?
- External stakeholder engagement
  - · Interest from Universities in the methodology for a teaching resource
  - A demonstration of practical geodesy techniques to generate interest of Geodesy as a career path for spatial science

Ē		On the derivation of parame	eters of
	$x_{IVI}$	radio telescopes for precise	
		determination of the invaria	ant
	Tzei		
		measurement point from a	point
m 1		cloud of measurements	
	Tzer	Bart Thomas, Jack McCubbine, Craig Harrison	
ose		Abstract	
ra	Wh	Abstract	
	$\Delta x_z$ $\Delta y_z$	This study develops a robust mathematical framework for precise	ly determining the
acc	$\Delta y_z$ $\Delta z_z$	invariant measurement point (IVP), and axis orientation of a "two	
n, pet	$\Delta z_{Z}$	Zenith) radio telescope from a point cloud generated through met	
tho	with	observations. Utilizing a series of transformations and rotations, v crucial for improving the accuracy of celestial measurements. Our	
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	Rod		
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iabi	SAZ	Mathematical model	
rep asu	- 12	The mathematical model for the position of a point $\mathbf{x}_{p_0} = [x_{p_0}, y_p]$	.z. ] on the dish
ults		following one rotation through the zenith axis and on rotation through	such the azimuth
		axis to give a point $\mathbf{x}_{p_1} = [\mathbf{x}_{p_1}, \mathbf{y}_{p_2}, \mathbf{z}_{p_1}]$ is given by equation 1.	
e re			
icie	so ti	$[\mathbf{x}_{ps}, 1]^T = T_{IVP} * R_A T_{Zen} * R_{Zen} T_{Zen} - T_{IVP} [\mathbf{x}_{ps}, 1]^T$	(1)
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ima	RAZ	Where,	
	RAT	$T_{IVP} = \begin{bmatrix} 1 & 0 & 0 & x_{Ivp} \\ 0 & 1 & 0 & y_{Ivp} \\ 0 & 0 & 1 & z_{Ivp} \end{bmatrix}$	(1 a)
or			
	Sim		
ne	rota		
110	mat	$T_{IVP^{*}} = \begin{bmatrix} 1 & 0 & 0 & -x_{top} \\ 0 & 1 & 0 & -y_{top} \\ 0 & 0 & 1 & -z_{top} \\ 0 & 0 & 0 & -z_{top} \end{bmatrix}$	
		$T_{IVP^{-}} = \begin{bmatrix} 0 & 1 & 0 & -y_{ivp} \\ 0 & 0 & 1 & -y_{ivp} \end{bmatrix}$	(1 b)
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# Thank You

#### Thanks also for the contribution and assistance of:

Craig Harrison, Renato Gonzalez, Jack McCubbine, Randall Carmen, Brett Reid (UTAS), Warren Hankey (UTAS), Boye Zhou (UTAS), Darryl Stacey (NTGS), Darryn Whelan (CDU), NT Survey Services, Jamie McCallum (UTAS), Lucia McCallum (UTAS), Christopher Watson (UTAS),

...who all helped make the latest survey campaign possible.

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Geoscience Australia acknowledges the Traditional Custodians of Country throughout Australia and recognises the continuing connection to lands, waters and communities. We pay our respects to Aboriginal and Torres Strait Islanders Cultures: and to elders past, present and emerging.

# The most relevant SDGs related to the presentation and theme of this session



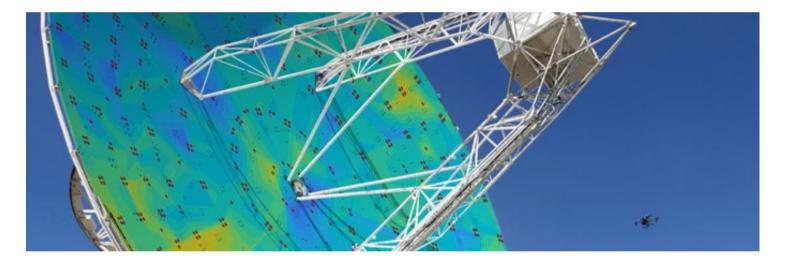
## SUSTAINABLE G ALS

International Federation of Surveyors supports the Sustainable Development Goals

## A quick shout-out to some novel research:

Research into effect of gravity deformation on VLBI antenna

• 26-m radio telescope at the Mount Pleasant Radio Observatory Hobart (Tasmania, Australia)<sup>1</sup>



<sup>1</sup>Lösler, M., Eschelbach, C., Greiwe, A., Zhou, B., McCallum, L. Innovative approach for modelling gravity-induced signal path variations of VLBI radio telescopes. *Earth Planets Space* **77**, 9 (2025). https://doi.org/10.1186/s40623-024-02110-8